

PHY5524 Problem Set 8: Solution

Problem 1

(a) The partition function for a single particle in a classical 3D gas is

$$Q_1^{(3D)} = \int \frac{d^3r d^3p}{h^3} e^{-\beta p^2/2m} = V \left(\frac{2\pi m k_B T}{h^2} \right)^{3/2} \quad (1)$$

For N particles the partition function is then

$$Q_N^{(3D)} = \frac{1}{N!} \left(Q_1^{(3D)} \right)^N \quad (2)$$

(b) To find the chemical potential we first find the Helmholtz free energy

$$A^{(3D)} = -k_B T \ln Q_N^{(3D)} = k_B T N \left(\ln \frac{N}{Q_1^{(3D)}} - 1 \right) \quad (3)$$

(where we have used Stirling's approximation $\ln N! = N \ln N - N$. The chemical potential is then

$$\mu_{3D} = \left(\frac{\partial A}{\partial N} \right)_{T,V} = k_B T \ln \frac{N}{Q_1^{(3D)}} \quad (4)$$

where $N = N_1$, the number of particles in the 3D gas.

(c) The partition function for a single particle in a classical 2D gas with energy $E = p^2/2m - \eta$ is

$$Q_1^{(2D)} = \int \frac{d^2r d^2p}{h^2} e^{-\beta(p^2/2m - \eta)} = A \left(\frac{2\pi m k_B T}{h^2} \right)^2 e^{\beta\eta} \quad (5)$$

(Here A is the area of the gas — not to be confused with the free energy!) For N particles the partition function is then

$$Q_N^{(2D)} = \frac{1}{N!} \left(Q_1^{(2D)} \right)^N \quad (6)$$

and the Helmholtz free energy is

$$A^{(2D)} = -k_B T \ln Q_N^{(2D)} = k_B T N \left(\ln \frac{N}{Q_1^{(2D)}} - 1 \right) \quad (7)$$

and, finally, the chemical potential of the 2D gas is

$$\mu_{2D} = \left(\frac{\partial A}{\partial N} \right)_{T,V} = k_B T \ln \frac{N}{Q_1^{(2D)}} \quad (8)$$

where $N = N_2$, the number of particles in the 2D gas.

(d) If the 2D and 3D gases are in equilibrium, their chemical potentials must be equal. Thus we require

$$\mu_{3D} = \mu_{2D} \quad (9)$$

which gives

$$k_B T \ln \frac{N_1}{Q_1^{(3D)}} = k_B T \ln \frac{N_2}{Q_1^{(2D)}} \quad (10)$$

This in turn implies

$$\frac{N_1}{Q_1^{(3D)}} = \frac{N_2}{Q_1^{(2D)}} \quad (11)$$

which, after plugging in expressions for $Q^{(3D)}$ and $Q^{(2D)}$, yields

$$\frac{N_2}{A} = \frac{N_1}{V} \left(\frac{h^2}{2mk_B T} \right)^{1/2} e^{\eta/k_B T} \quad (12)$$

Finally, we can compute the pressure of the 3D gas as follows

$$P = - \left(\frac{\partial A^{(3D)}}{\partial V} \right)_{N,T} = N_1 \frac{k_B T}{V} \quad (13)$$

This, of course, implies that the 3D gas satisfies the ideal gas law. Thus we have

$$\frac{N_1}{V} = \frac{P}{k_B T} \quad (14)$$

Putting everything together we obtain the following expression for the density of the 2D gas as a function of the pressure of the 3D gas,

$$n = \frac{N_2}{A} = \frac{P}{k_B T} \left(\frac{h^2}{2mk_B T} \right)^{1/2} e^{\eta/k_B T} \quad (15)$$

Problem 2

(a) For a 2D gas with dispersion $\epsilon(\vec{k}) = \hbar^2 k^2 / 2m$, the density of states is given by

$$a(\mathcal{E}) = A \int \frac{d^2 k}{(2\pi)^2} \delta \left(\mathcal{E} - \frac{\hbar^2 k^2}{2m} \right) \quad (16)$$

where A is the area of the 2D gas. For $\mathcal{E} < 0$ the delta function is never satisfied and $a(\mathcal{E})$ is zero. For $\mathcal{E} > 0$ we evaluate $a(\mathcal{E})$ as follows

$$a(\mathcal{E}) = \frac{A}{(2\pi)^2} 2\pi \int_0^\infty k dk \frac{\delta(k - \sqrt{2E/m})}{|\hbar^2 k/m|} \quad (17)$$

$$= A \frac{m}{2\pi \hbar^2} \quad (18)$$

Thus we see that in 2D the density of states is constant for $\mathcal{E} > 0$ and zero for $\mathcal{E} < 0$.

(b) Expressing the density of bosons as a sum over Bose occupation factors we have

$$\frac{N}{A} = \frac{1}{A} \sum_k n(\mathcal{E}(k)) = \frac{1}{A} n(0) + \int \frac{d^2 k}{(2\pi)^2} \frac{1}{z_1 e^{\beta \hbar^2 k^2 / 2m} - 1} \quad (19)$$

where, in the second equality, we have separated out the occupation of the ground state $n(0)$ and converted the remaining sum over \vec{k} to an integral.

Note that

$$n(0) = \frac{1}{z^{-1} - 1} = \frac{z}{1 - z} \quad (20)$$

If we substitute the above in for $n(0)$ and, using the density of states from (a) convert the integral over k into an integral over \mathcal{E} , we have

$$\frac{N}{A} = \frac{z}{1 - z} \frac{1}{A} + \frac{2\pi m}{h^2} \int_0^\infty d\mathcal{E} \frac{1}{z^{-1} e^{\beta \mathcal{E}} - 1} \quad (21)$$

Finally, making the change of variables $y = \beta \mathcal{E} = \mathcal{E} / (k_B T)$ in the integral, we obtain the desired result

$$\frac{N}{A} = \frac{z}{1 - z} \frac{1}{A} + \frac{2\pi k_B T m}{h^2} \int_0^\infty dy \frac{1}{z^{-1} e^y - 1} \quad (22)$$

(c) The integral appearing in (22) can be done analytically with the result

$$\int_0^\infty dy \frac{1}{z^{-1} e^y - 1} = \ln \frac{1}{1 - z} \quad (23)$$

Thus we have

$$\frac{N}{A} = \frac{N_0}{A} - \frac{1}{\lambda^2} \ln(1-z) \quad (24)$$

In the thermodynamic limit, at $T = T_c$ the condensate density N_0/A vanishes, and the chemical potential μ goes to zero so the fugacity z goes to 1. Because the integral (23) diverges as $z \rightarrow 1$, it follows that $T_c = 0$ for this gas.

(d) Because this 2D gas does not exhibit Bose condensation, at any finite temperature in the thermodynamic limit the condensate density vanishes. Thus we have

$$\frac{N}{A} = \frac{1}{l^2} = \frac{1}{\lambda^2} \ln \frac{1}{1-z} \quad (25)$$

Here $l = (A/N)^{1/2}$ is the mean interparticle spacing, and $\lambda = \left(\frac{\hbar^2}{2\pi m k_B T}\right)^{1/2}$ is the thermal wavelength. Solving for the fugacity z we obtain

$$z = 1 - e^{-\lambda^2/l^2} \quad (26)$$

In the classical limit, i.e. the high temperature / low density limit, $\lambda^2/l^2 \ll 1$. Thus in this limit

$$z \simeq 1 - \left(1 - \frac{\lambda^2}{l^2}\right) = \frac{\lambda^2}{l^2} \ll 1 \quad (27)$$

as expected.

Problem 3

(a) As in the previous problem, we use the fact that in the thermodynamic limit, at T_c the fugacity is 1 ($z = 1$) and the condensate density vanishes ($N_0/A = 0$) to write down the following equation which we can solve to find T_c

$$\frac{N}{A} = \int \frac{d^2k}{(2\pi)^2} \frac{1}{e^{\beta c(\hbar k)^{3/2}} - 1} \quad (28)$$

$$= \int \frac{d^2k}{(2\pi)^2} \frac{1}{e^{\alpha k^{3/2}/T} - 1} \quad (29)$$

where $\alpha = c\hbar^{3/2}/k_B$.

Making the change of variables $\vec{q} = \vec{k}/T^{2/3}$ we find

$$\frac{N}{A} = T_c^{4/3} \int \frac{d^2q}{(2\pi)^2} \frac{1}{e^{\alpha q^{3/2}} - 1} \quad (30)$$

For T_c to be finite it is only necessary that the integral over q appearing in the above expression be finite. The only ‘‘dangerous’’ limit of the integrand is for small q . In this limit we can Taylor expand $e^{\alpha q^{3/2}}$ to find that

$$\int \frac{d^2q}{(2\pi)^2} \frac{1}{e^{\alpha q^{3/2}} - 1} = \int_0^\infty q dq \frac{1}{q^{3/2}} \int_0^{2\pi} \frac{dq}{q^{1/2}} = \text{finite}. \quad (31)$$

The point is that the small q singularity in the integrand goes as $1/q^{1/2}$, which is integrable. (This is in contrast to the previous problem, for which the same analysis would give a singularity of the form $1/q$ which is *not* integrable.) Because this integral is finite we conclude that T_c is finite.

Given that T_c is finite, we can also conclude from the above that

$$T_c \propto n^{3/4} \quad (32)$$

where $n = N/A$ is the number density of bosons.

(b) The grand partition function for bosons with dispersion $\mathcal{E}(\vec{k})$ is

$$\mathcal{Z} = \prod_{\vec{k}} \frac{1}{1 - ze^{-\beta \mathcal{E}(\vec{k})}} \quad (33)$$

The grand potential is then

$$\Sigma = -k_B T \ln \mathcal{Z} = k_B T \sum_{\vec{k}} \ln(1 - ze^{-\beta \mathcal{E}(\vec{k})}) \quad (34)$$

Separating out the contribution of the condensate and turning the sum into an integral in the usual way we obtain

$$\Sigma = -k_B T \ln(1 - z) - k_B T A \int \frac{d^2 k}{(2\pi)^2} \ln(1 - ze^{-\beta \mathcal{E}(\vec{k})}) \quad (35)$$

Note that

$$N_0 + 1 = \frac{z}{1 - z} + 1 = \frac{1}{1 - z} \quad (36)$$

where $N_0 = n(0)$ is the occupation number of the ground state. It follows that

$$\Sigma = k_B T \ln(N_0 + 1) - k_B T A \int \frac{d^2 k}{(2\pi)^2} \ln(1 - ze^{-\beta \mathcal{E}(\vec{k})}) \quad (37)$$

The first term is $\propto \ln N_0$, while the second term is $\propto A$, where A is the area of the system. In the thermodynamic limit the first term, which is not extensive, is negligible when compared with the second term which *is* extensive. For $T < T_C$, again in the thermodynamic limit, the fugacity z is equal to 1. Thus, for $T < T_C$, in the thermodynamic limit the grand potential is given by

$$\Sigma = -k_B T \int \frac{d^2 k}{(2\pi)^2} \ln(1 - e^{-\alpha k^3/2/T}) \quad (38)$$

$$(39)$$

where α is as defined in (a) above.

To analyze Σ without actually doing the integral, we can again make the change of variables $\vec{q} = \vec{k}/T^{2/3}$, with the result

$$\Sigma = -(k_B T) T^{4/3} A \int \frac{d^2 q}{(2\pi)^2} \ln(1 - e^{-\alpha q^3/2}) \propto AT^{7/3} \quad (40)$$

Thus we see that for $T < T_C$, $\Sigma \propto T^{7/3}$. It follows that the entropy scales as

$$S = - \left(\frac{\partial \Sigma}{\partial T} \right)_{\mu, A} \propto T^{4/3} \quad (41)$$

and the pressure scales as

$$P = - \frac{\Sigma}{A} \propto T^{7/3} \quad (42)$$